

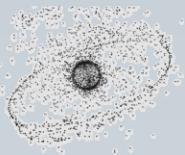
# Long-term Evolution of High Area-to-Mass Ratio Objects in Different Orbital Regions

T. Schildknecht, A. Vananti, A. Hinze

*Astronomical Institute, University of Bern, Switzerland*

PEDAS1-0017-12

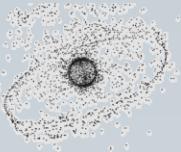
39th COSPAR Scientific Assembly, 14 – 22 July, 2012, Mysore,  
India



# Outline

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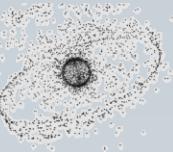
1. The AIUB/ESA HAMR Catalogue
2. Orbit Evolution in GEO & MEO
3. Summary



# Catalogue of Small-Size Space Debris

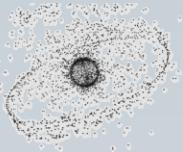
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- Build-up and maintenance of orbit catalogue of decimeter-sized debris in GEO/GTO/MEO
- Why?
  - Density/collision risk lower than in LEO  
**BUT:**  
**No sinks → population constantly grows**  
**→ Mitigation of debris is important**
- Need to know nature and sources of debris
  - Requires:
    - Orbit catalogue
    - Constant monitoring due to perturbations by non-gravitational forces



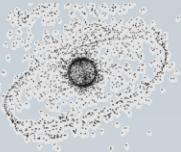
# Sensors





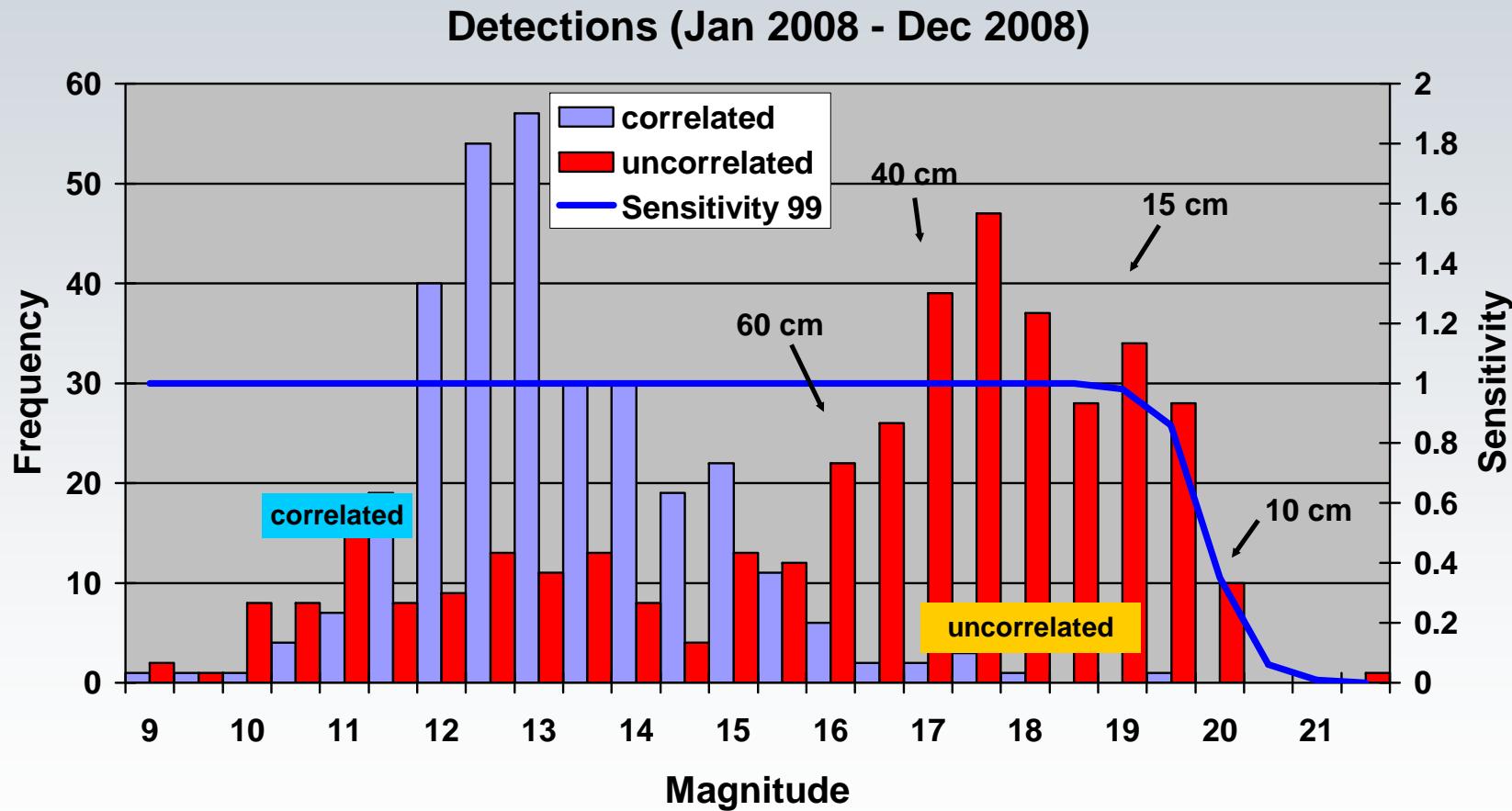
# High AMR GEO/GTO Catalogue

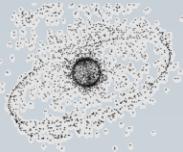
- **Discover new objects:** Obs. From Tenerife (OGS, AIUB)
  - **Secure orbits:** obs. from OGS, Zimmerwald (AIUB)
  - **Maintain orbits:** obs. from OGS, Zimmerwald, international partners, International Scientific Optical observation Network (ISON), ...
    - **Daily orbit maintenance** at AIUB and Keldysh Institute of Applied Mathematics of the Russian Academy of Sciences (KIAM)
- Orbit catalogue of high-altitude space debris
- **Provide predictions:**
    - To other partners (CNES, JAXA, NASA, Roscosmos...)
    - → to investigate physical properties of objects



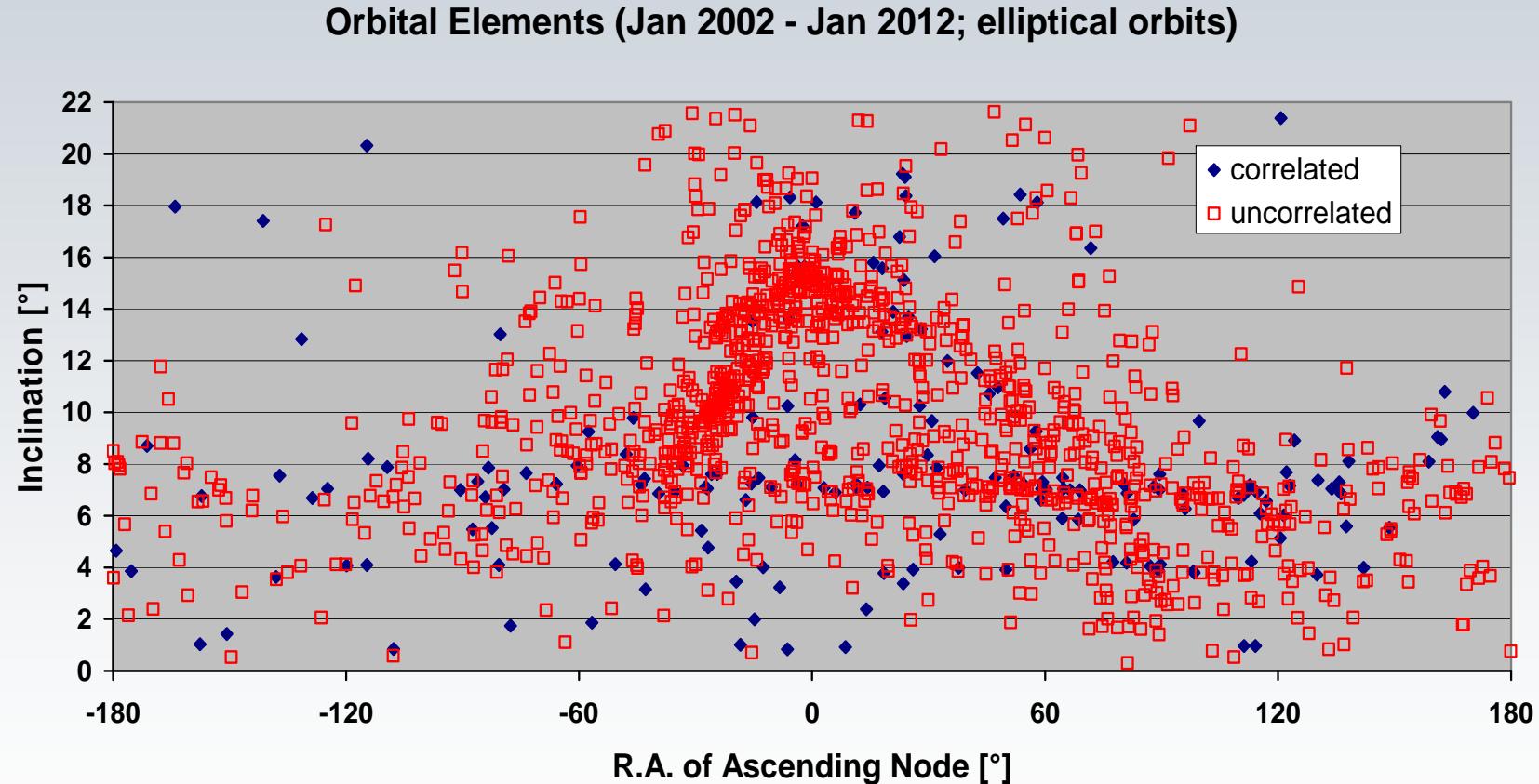
# ESA 2008 GEO/GTO Surveys

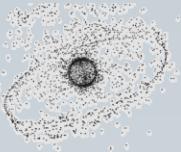
Continuous program, ~80 nights per year



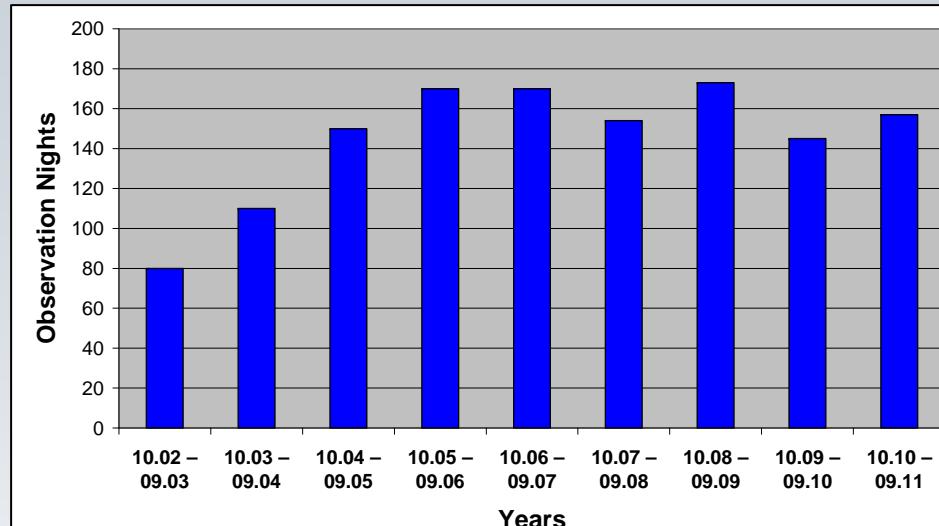


# 6-param. Orbits – $i$ vs $\Omega$

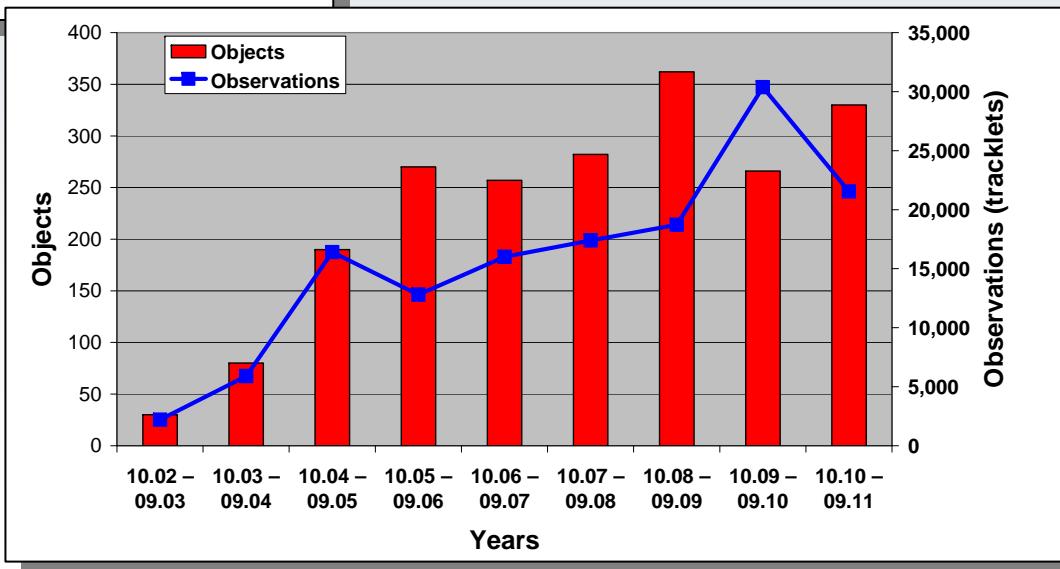




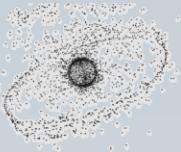
# “Routine” ZIMLAT Support



**ZIMLAT**  
Observation Nights

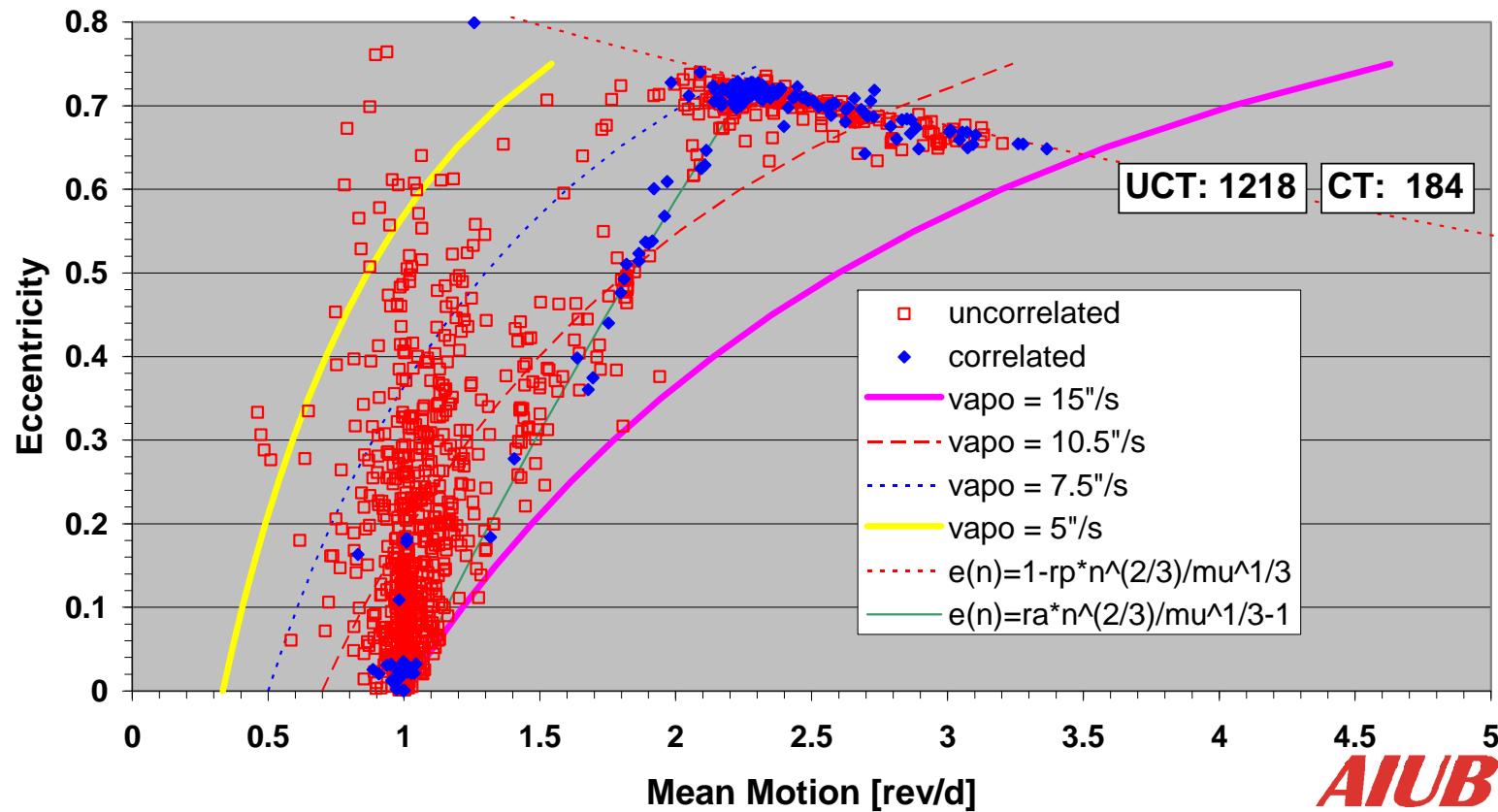


**ZIMLAT**  
Observations / Objects

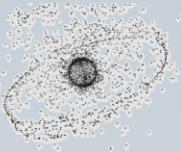


# High AMR GEO/GTO Catalogue

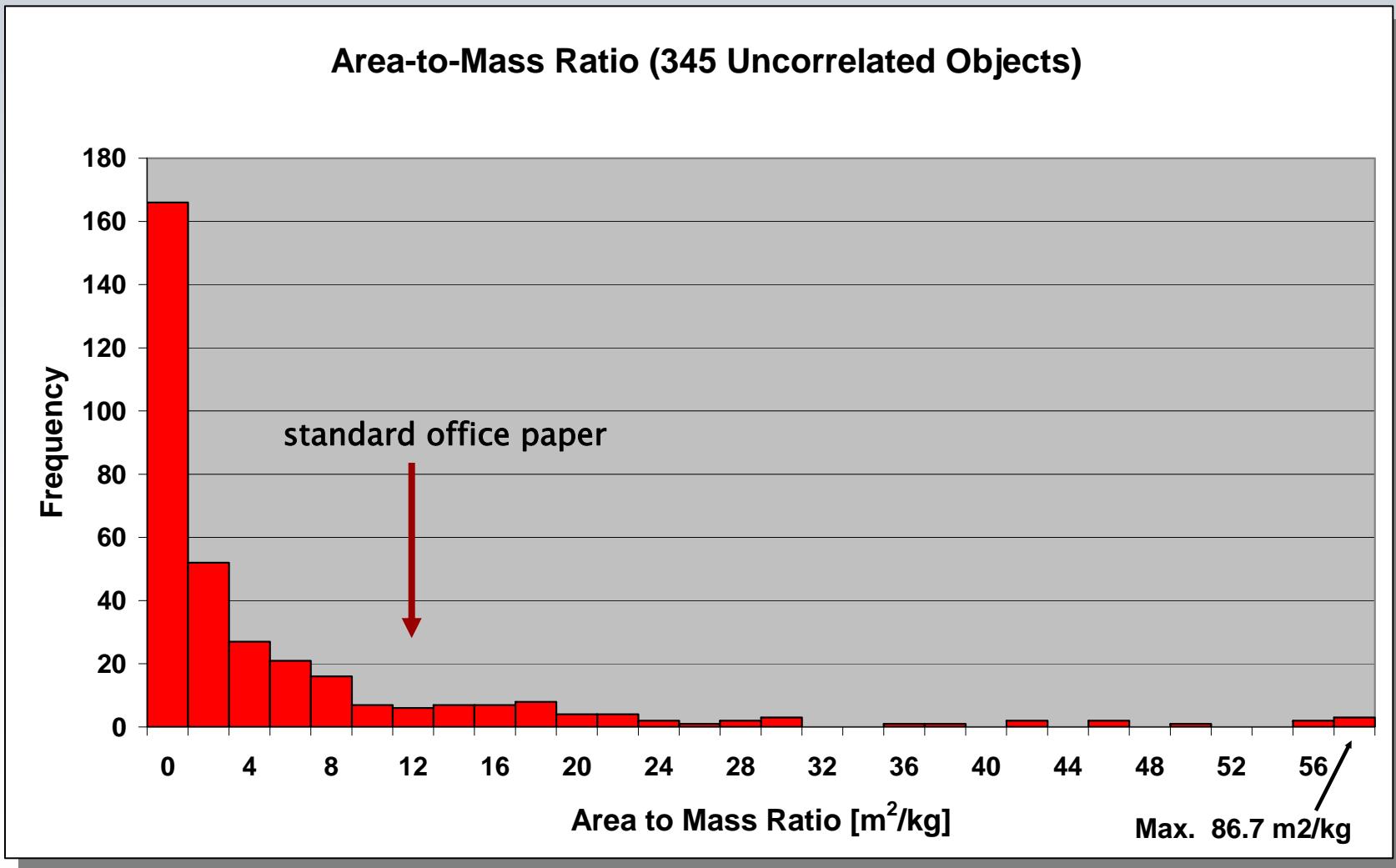
Eccentricity vs Mean Motion (Jan 2002 - Jan 2012; elliptical orbits)

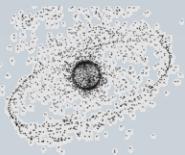


**AIUB**



# High AMR GEO/GTO Catalogue

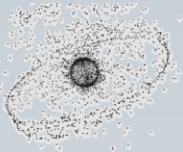




# Outline

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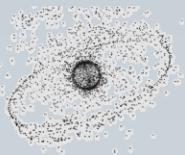
- 1. The AIUB/ESA HAMR Catalogue**
- 2. Orbit Evolution in GEO & MEO**
- 3. Summary**



# Methodology

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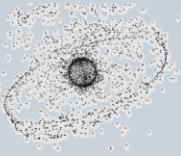
- **Assumption**
  - There are sources of high AMR debris when objects are left in space for long time span (aging processes, breakup-events, ...)
- **Question**
  - What are the impacts for the concept of "graveyard" orbits
- **Method**
  - Propagate orbits of a sample of observed debris over a time interval of 50 years assuming different AMR values (0.02, 0.02, 1, 5, 15 m<sup>2</sup>/kg).



# Propagator

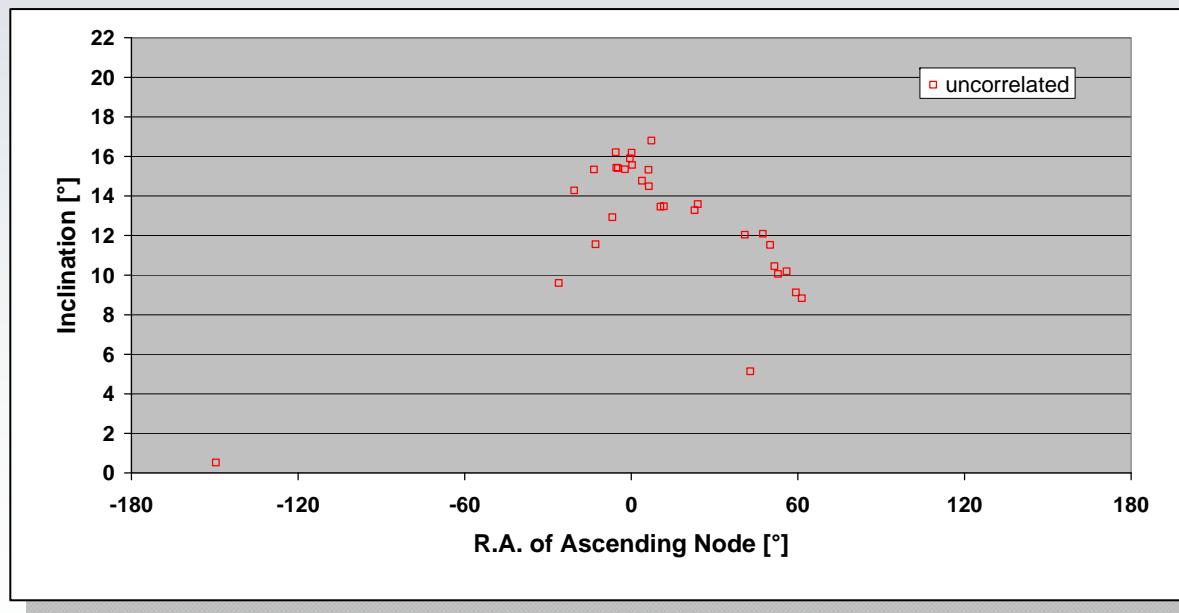
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- CelMech SATORB Propagator
  - numerical integrator (variable step size)
  - 12 x 12 Earth gravity field
  - gravitational perturbations from
    - Sun
    - Moon
    - Earth tides
  - corrections due to general relativity
  - direct radiation pressure (Sun only)
  - eclipses (Earth, Moon)

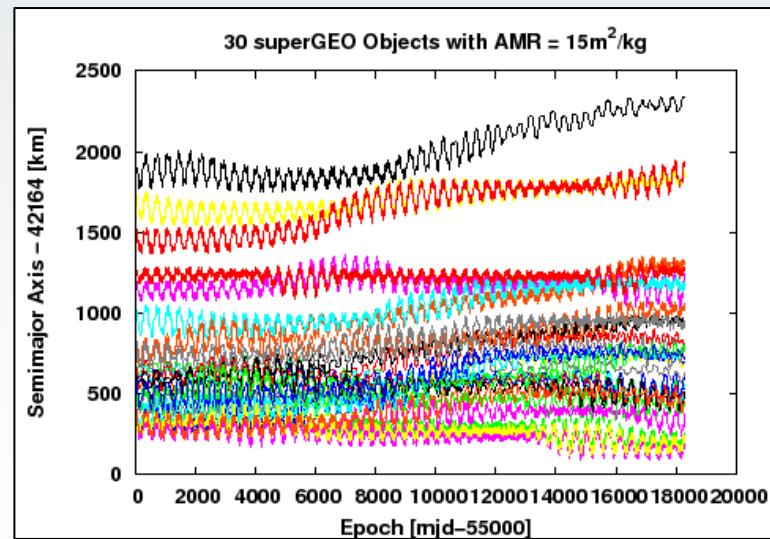
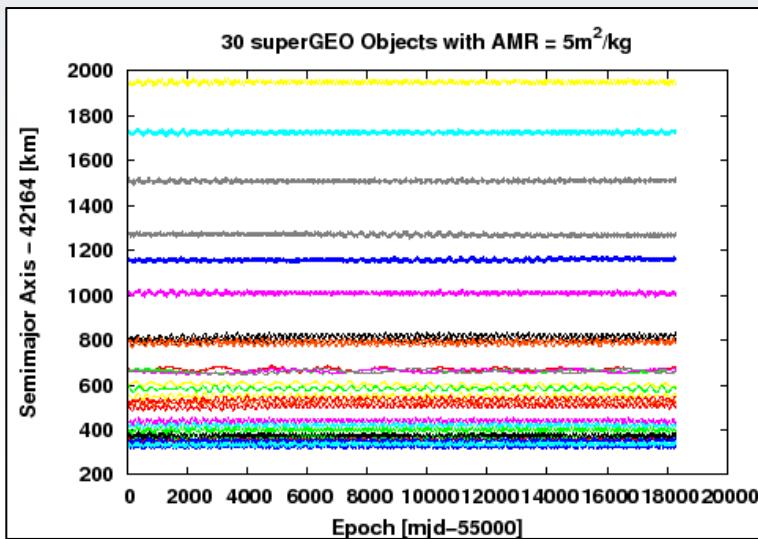
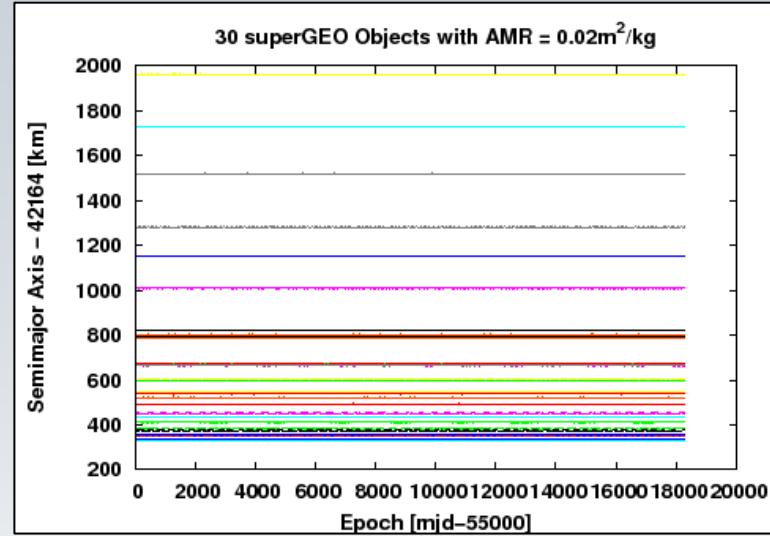
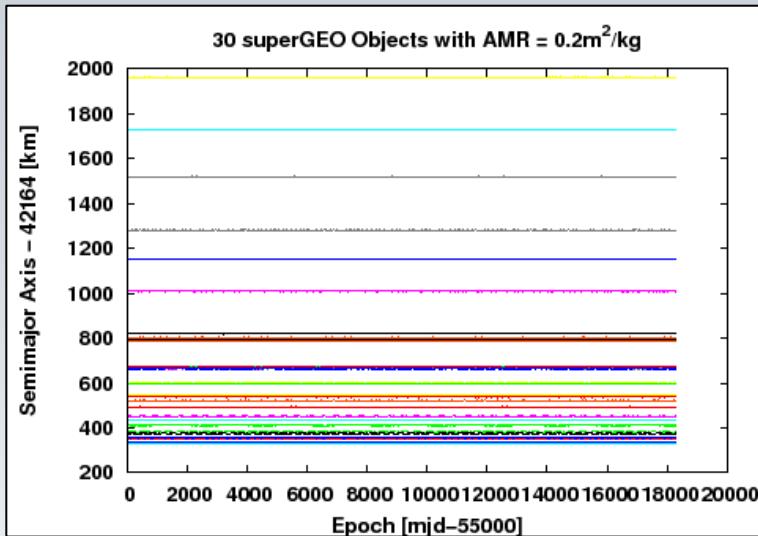


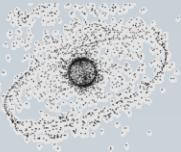
# Sample of super-GEO Objects

- Sample of 30 debris objects in super-GEO region
  - $a > 42464\text{km}$  ( $>300\text{km}$  above GEO)
  - $e < 0.05$ ; then set to 0.001
  - $t_0 = 55000\text{mjd}$  (2009-06-18)

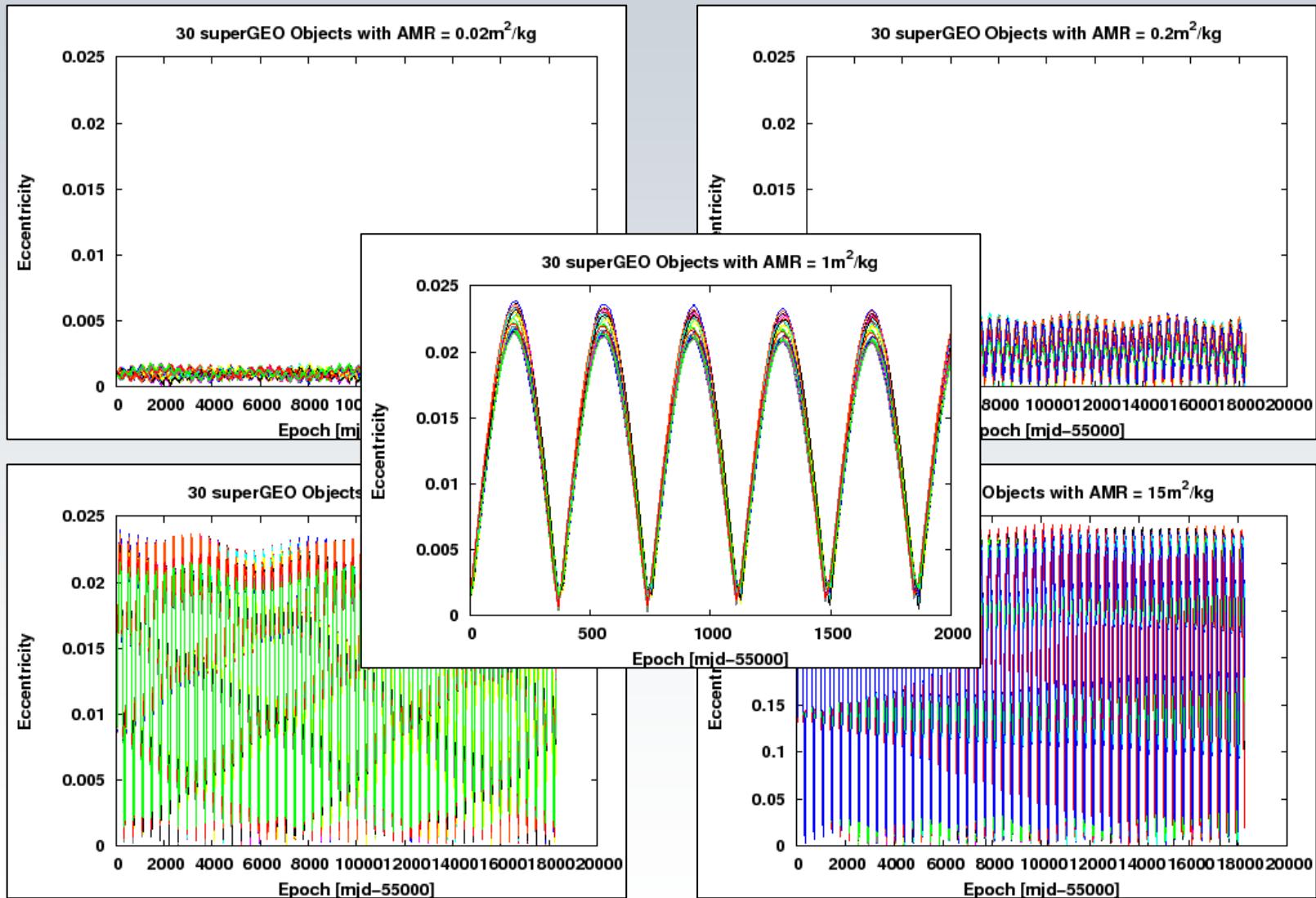


# Evolution of Semimajor Axis

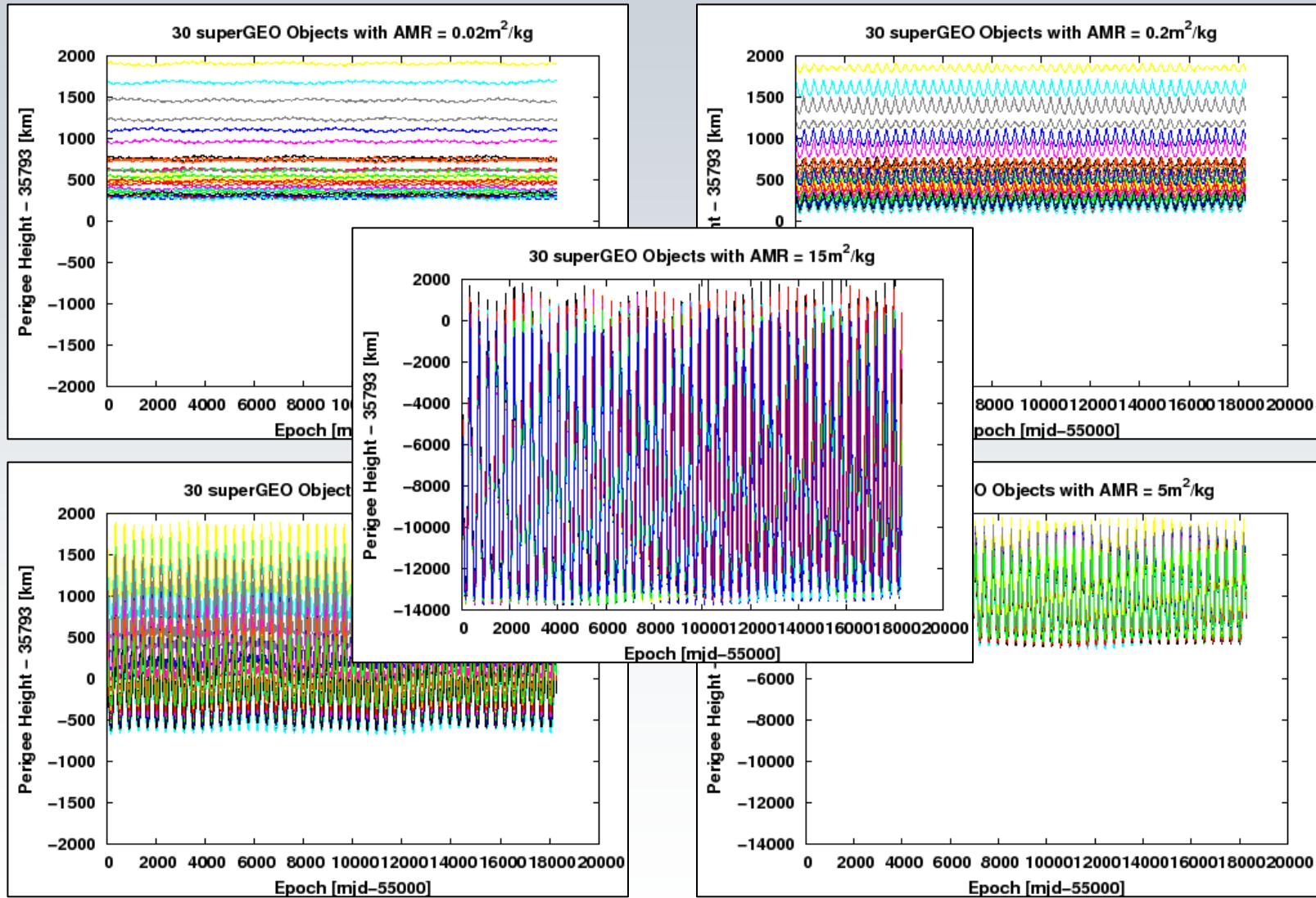


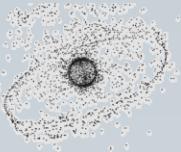


# Evolution of Eccentricity

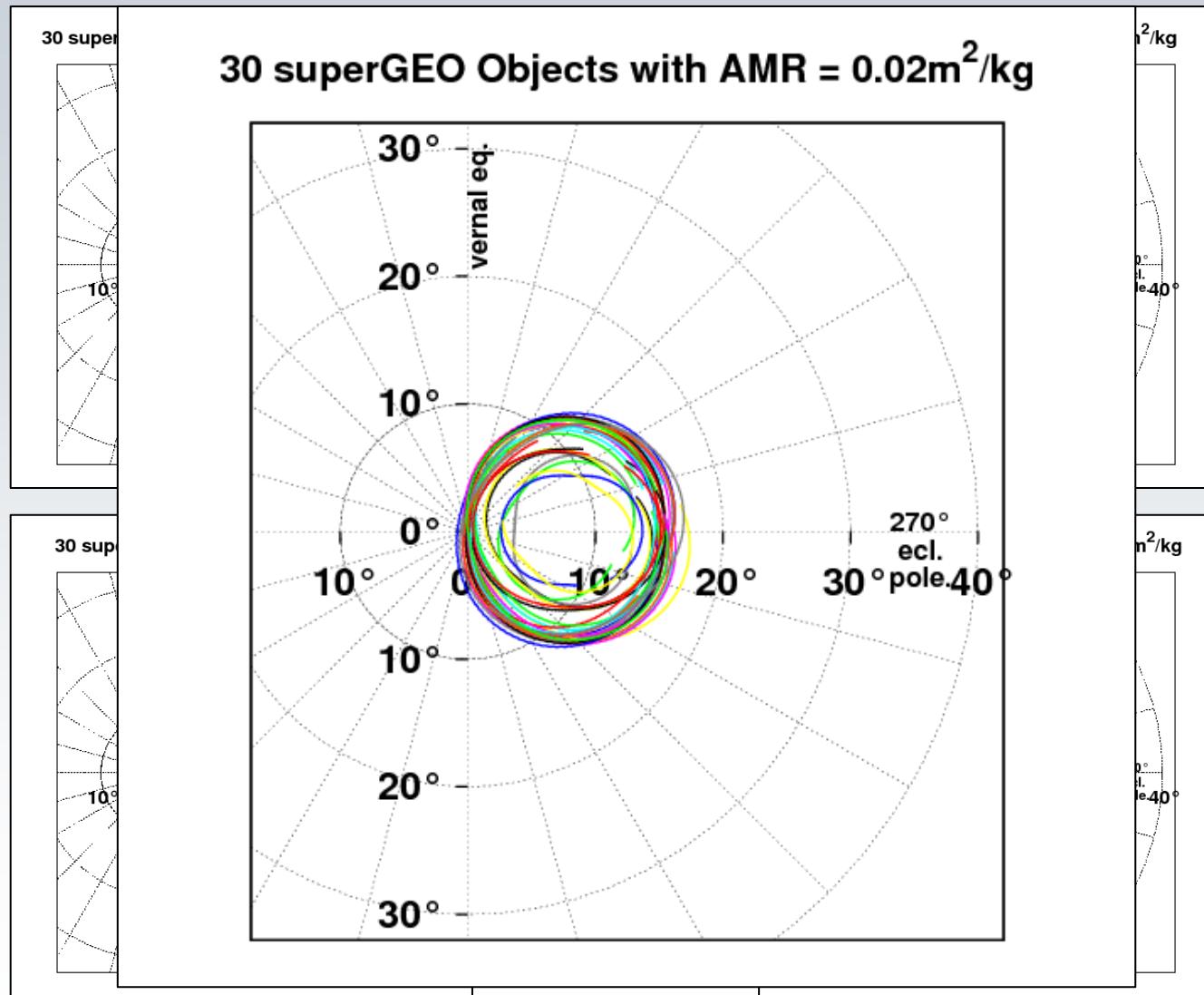


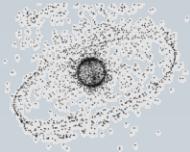
# Evolution of Perigee Height





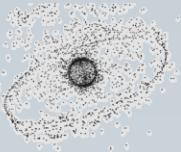
# Evolution of Orbital Plane





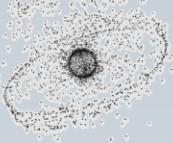
# GEO Graveyard

- Objects in GEO graveyard orbits ( $>300\text{km}$  above GEO) will cross GEO altitude if  $\text{AMR} > 0.2\text{m}^2/\text{kg}$ 
  - will also cross  $0^\circ$  GEO region for particular inclination/argument of perigee combinations
  - orbital plane and argument of perigee are both changing over time
  - precession of orbital planes not significantly changed for  $\text{AMR} < 5\text{m}^2/\text{kg}$
- Consistent with IADC deorbit guideline:  
$$\Delta H_{\min} = 235 + 1000 * \text{cr} * \text{AMR} [\text{km}]$$
  
→  $435\text{km}$  for  $\text{AMR} = 0.2\text{m}^2/\text{kg}$

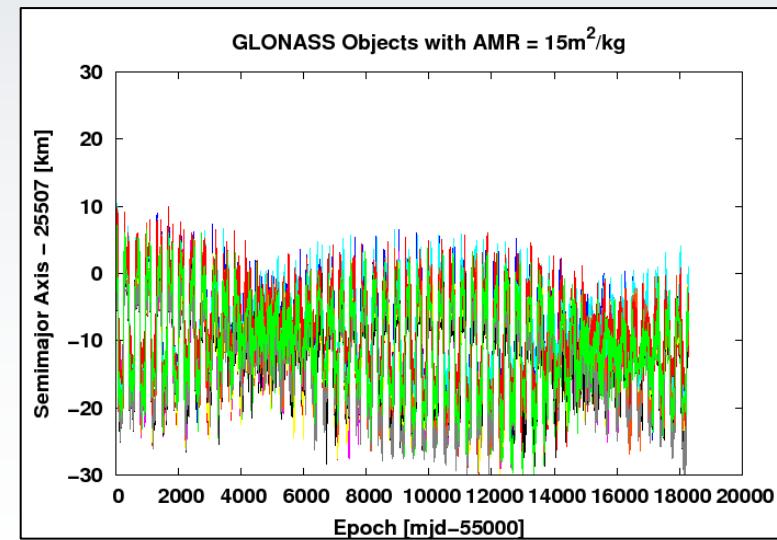
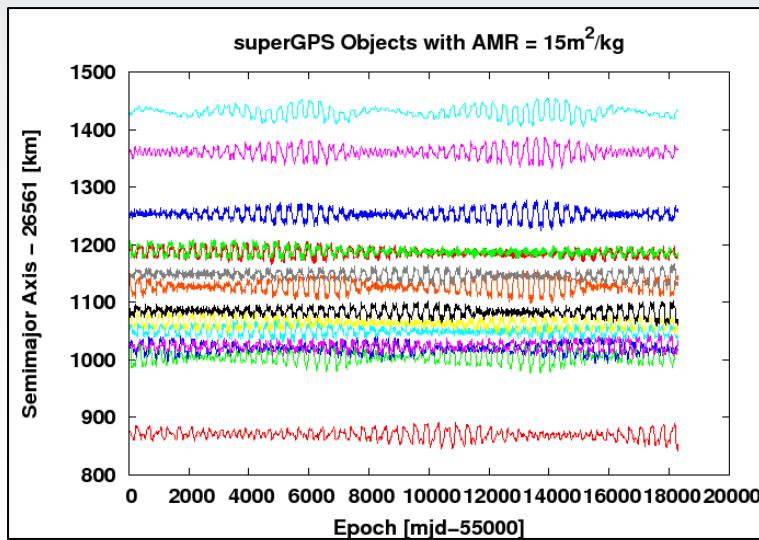
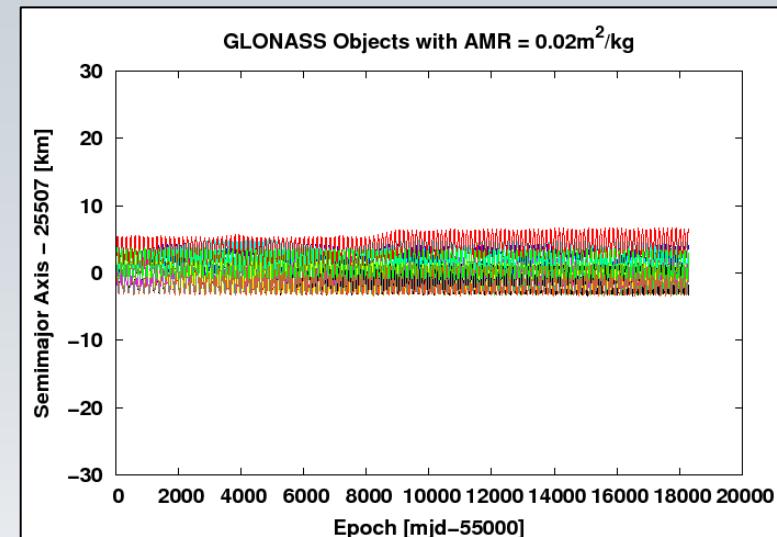
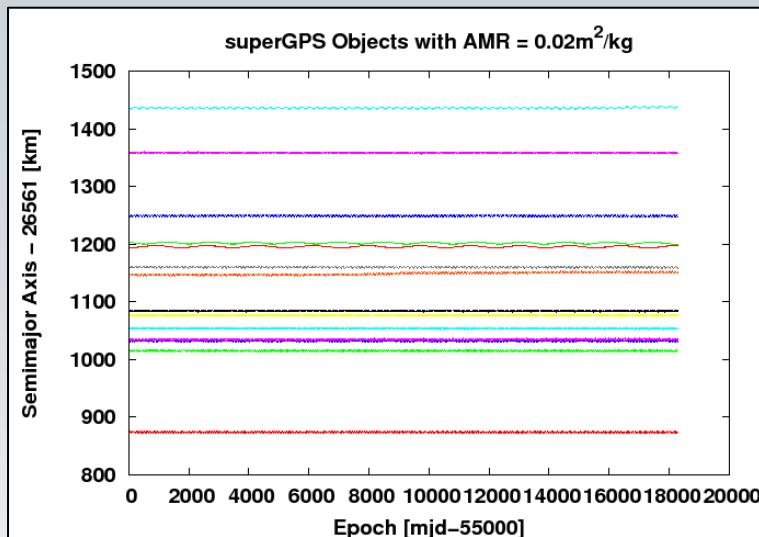


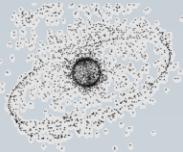
# Sample of super-GPS/GLONASS Objects

- Sample of super-GPS objects (14)
  - $a > 27400\text{km}$  ( $>900\text{km}$  above GPS)
  - all orbital elements at actual values
  - $t_0 = 55000\text{mjd}$  (2009-06-18)
- Sample of GLONASS objects in plane G2 (11)
  - $a \approx 25508\text{km}$
  - all orbital elements at actual values
  - $t_0 = 55000\text{mjd}$  (2009-06-18)

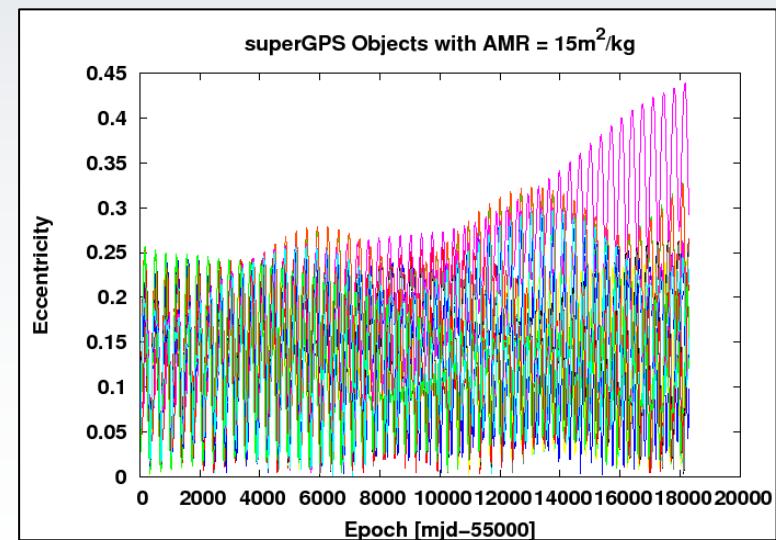
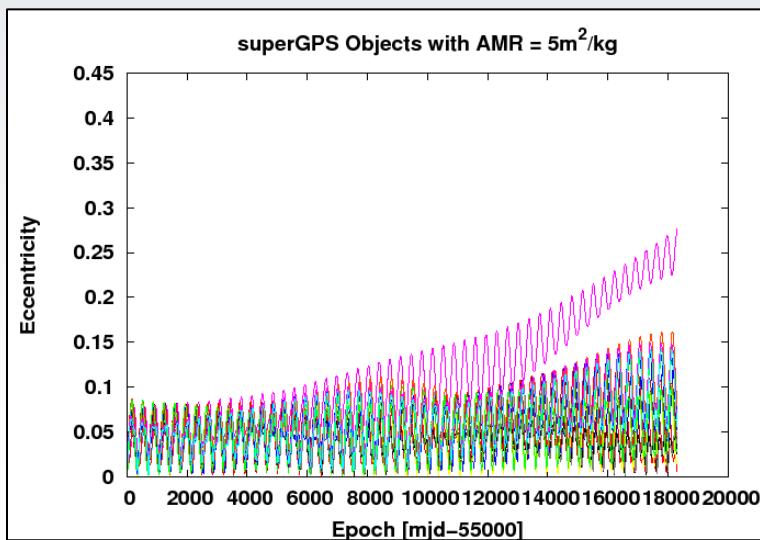
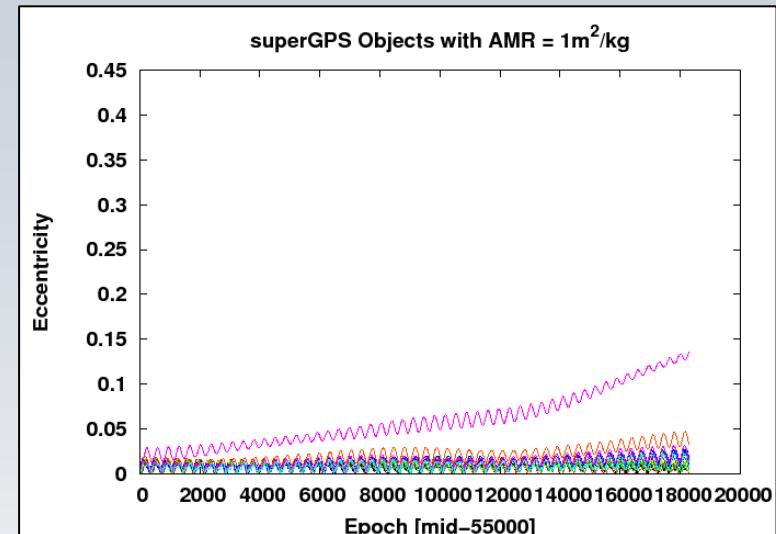
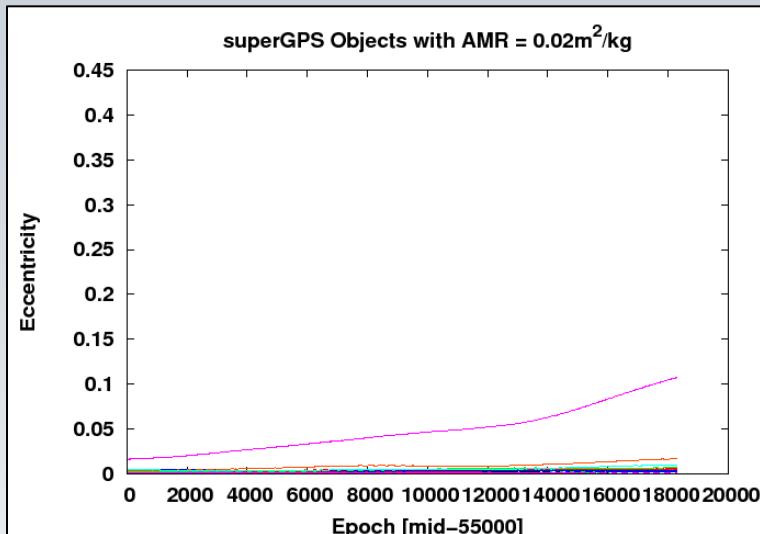


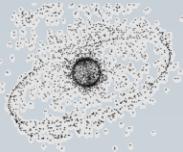
# Evolution of Semimajor Axis



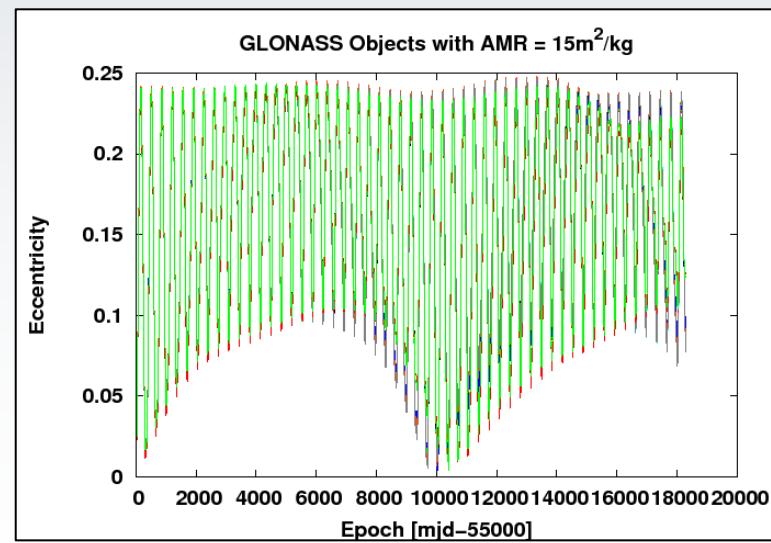
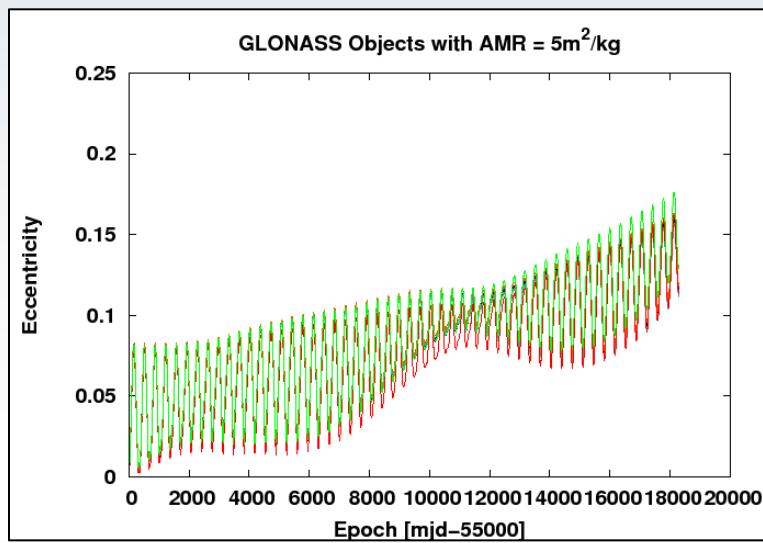
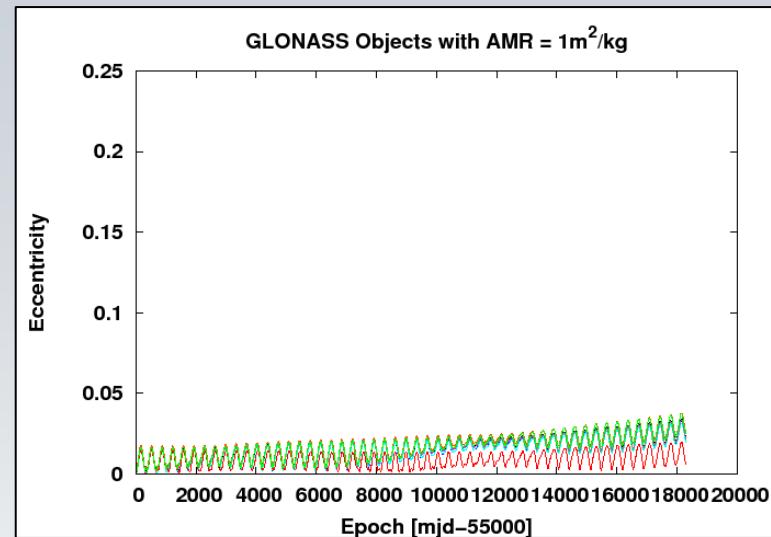
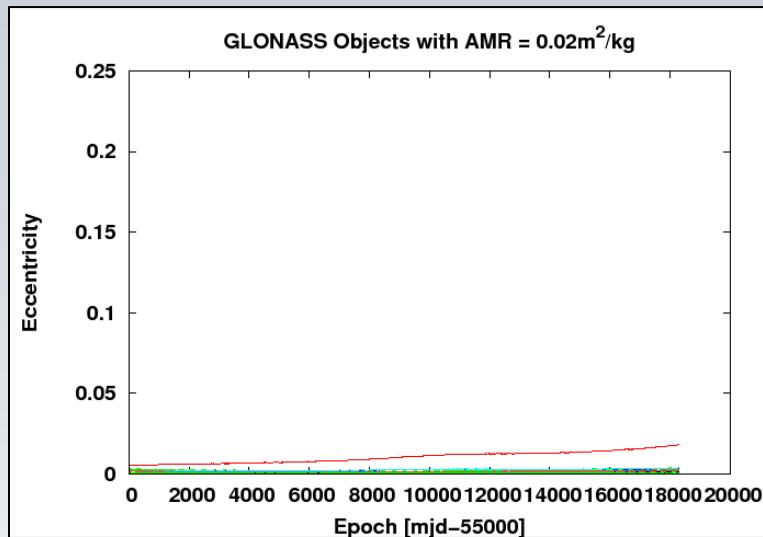


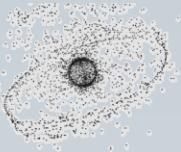
# Evolution of Eccentricity (GPS)



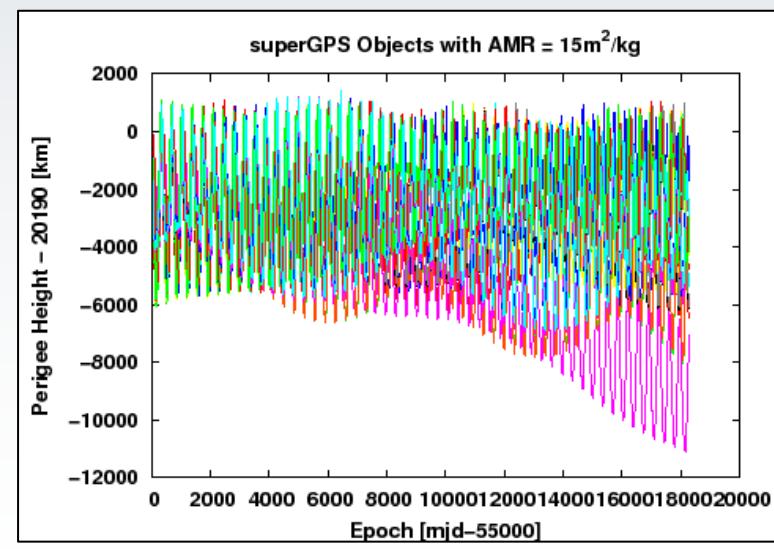
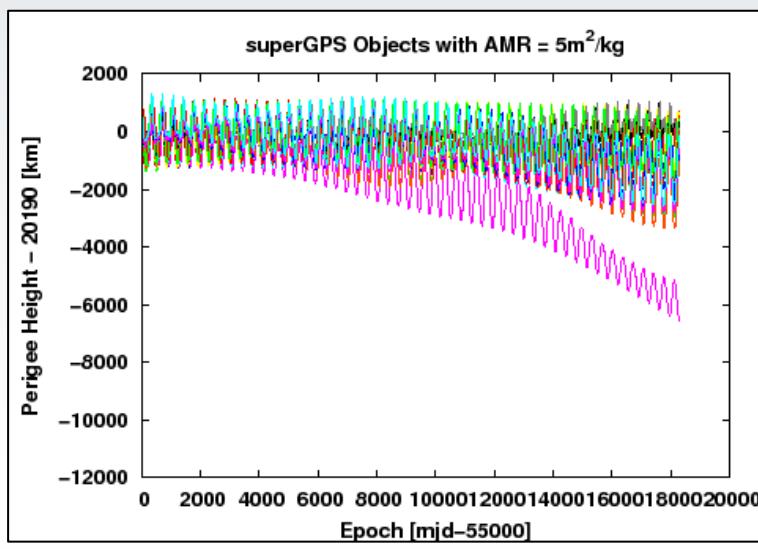
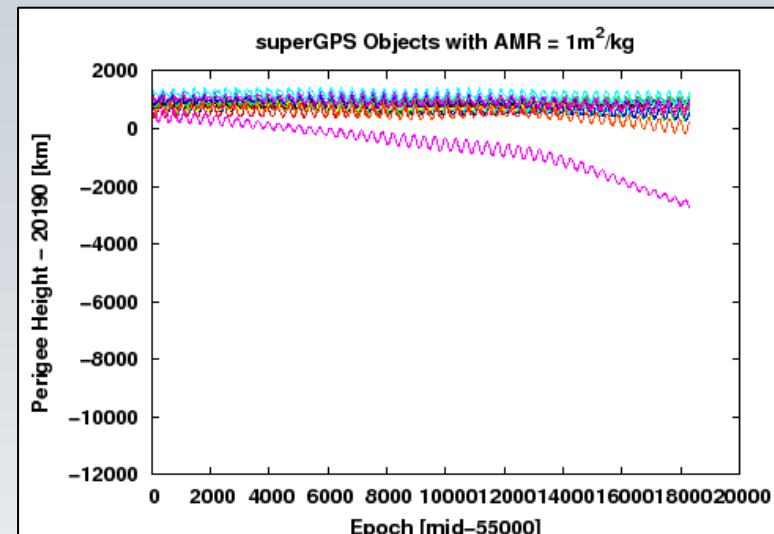
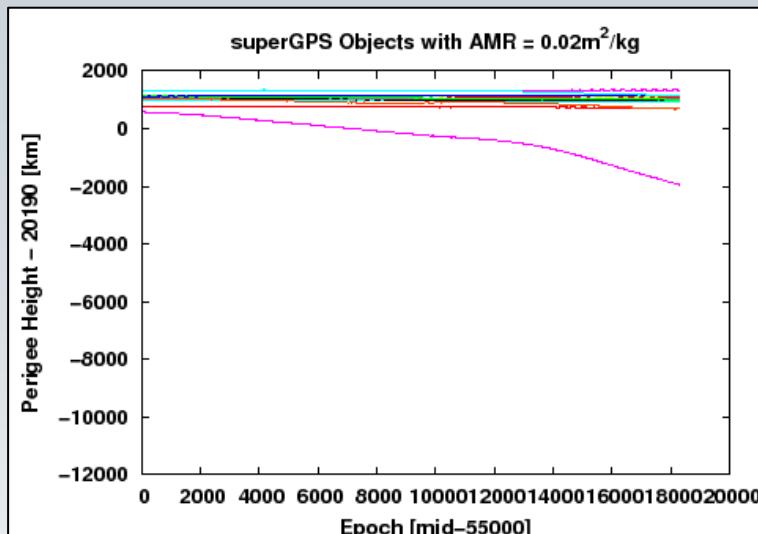


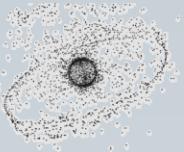
# Evolution of Eccentricity (GLONASS)



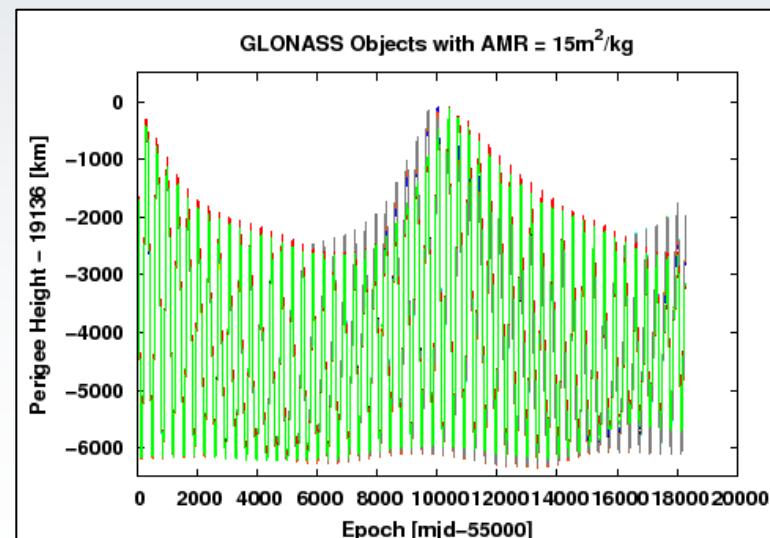
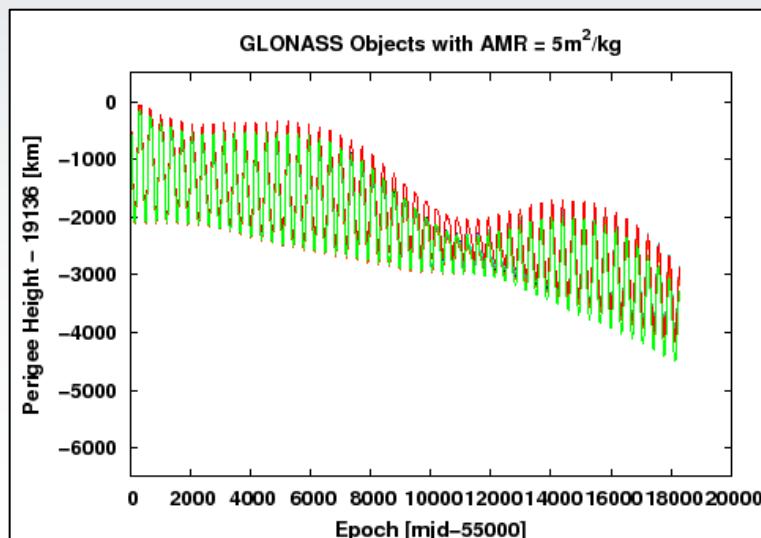
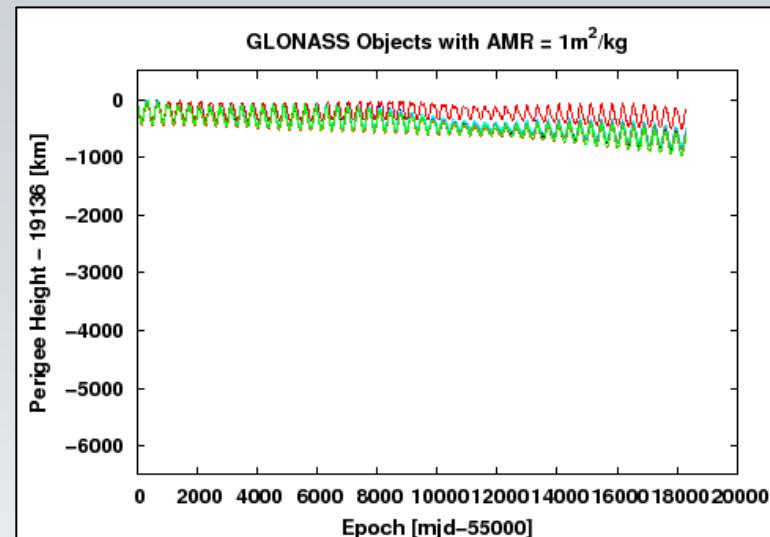
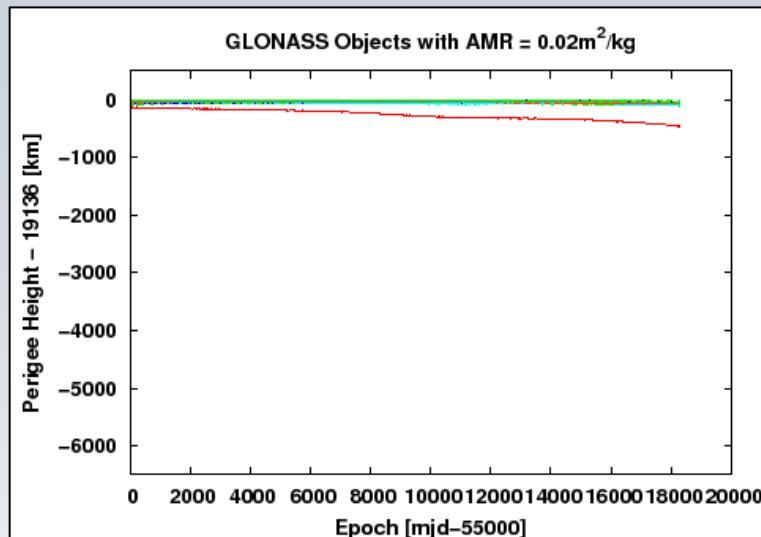


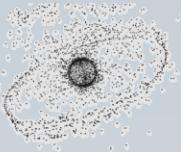
# Evolution of Perigee Height (GPS)



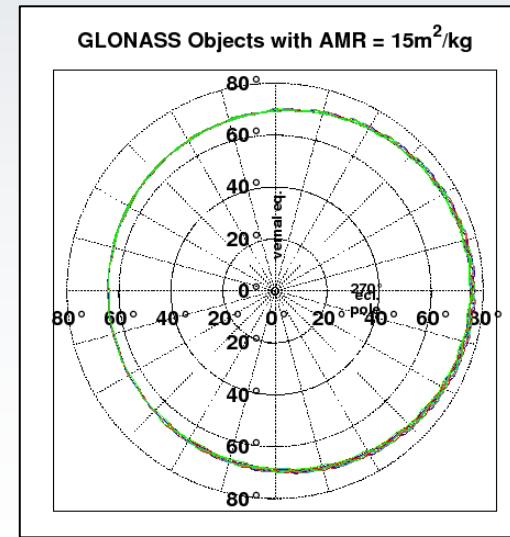
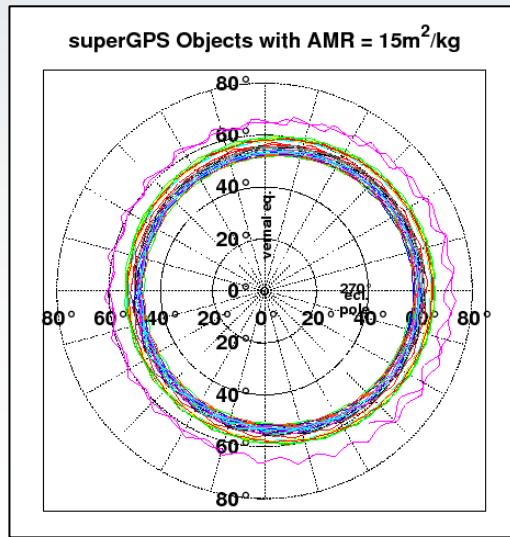
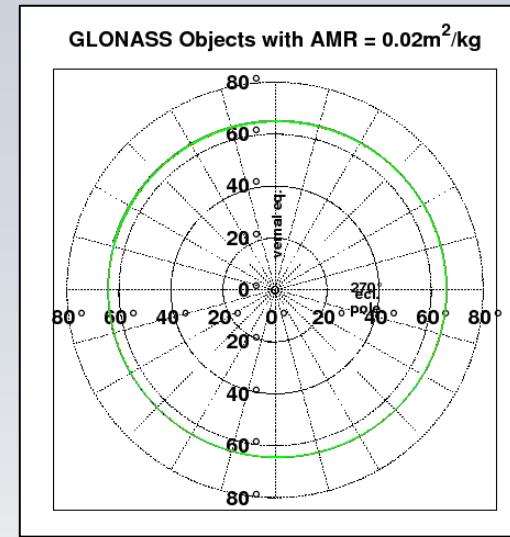
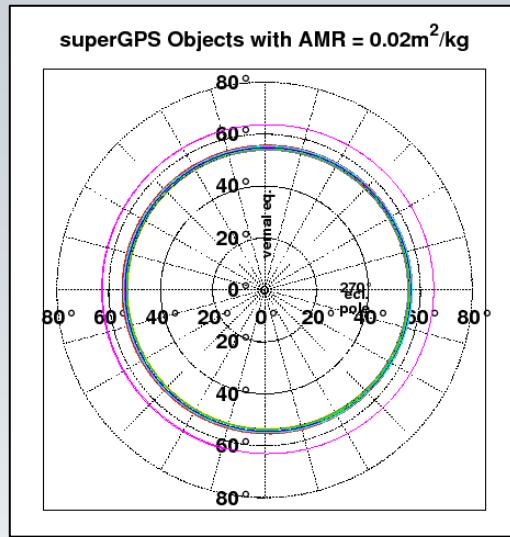


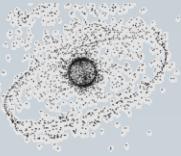
# Evolution of Perigee Height (GLONASS)





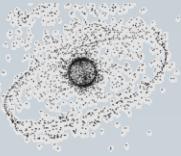
# Evolution of Orbital Plane





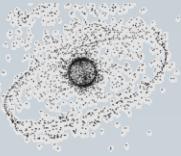
# GPS/GLONASS Orbits

- Strong gravitational perturbations of GPS orbits due to 2:1 mean motion resonance
  - eccentricity
  - inclination, precession of orbital plane
- Objects in GPS graveyard orbits ( $>900\text{km}$  above GPS) will cross GPS altitude if
  - $\text{AMR} \gtrsim \text{m}^2/\text{kg}$
  - OR
  - $e(t_0) \gtrsim 0.01$
- Perigee height of GLONASS orbits will change by  $>1000\text{km}$  if
  - $\text{AMR} \gtrsim 1\text{ m}^2/\text{kg}$



# Conclusions

- Catalogue of high AMR GEO/GTO objects
  - significant debris population with high AMR found in GEO and GTO region
  - orbits maintained by sharing the data in a network of observatories (KIAM, ISON)
- GEO/MEO graveyards
  - sources of high AMR debris to be expected when objects are left in space for long time span (aging processes, breakup-events, ...)
  - objects in GEO graveyard (300km above GEO) will cross GEO altitude if  $\text{AMR} > 0.2 \text{m}^2/\text{kg}$
  - perigee height of objects in GPS/GLONASS orbits will change for  $> 1000\text{km}$  if  $\text{AMR} \gtrsim 1 \text{m}^2/\text{kg}$
  - long-term evolution of GPS orbits dominated by gravitational resonance effects for  $e(t_0) \gg 0.01$



# Acknowledgments

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- Great thanks to our staff and observers at the OGS and Zimmerwald observatories!
- Support in the form observations to maintain the orbits is provided by the Keldysh Institute of Applied Mathematics (KIAM) in the framework of the ISON collaboration (AIUB–KIAM collaboration).